HSC Physics - Module 5: Advanced Mechanics - Motion in Gravitational Fields Short Answer Question Solutions 5
41.

| Criteria | Mark |
| :---: | :--- |
| $\bullet$ Correct substitution into a correct formula, yielding a correct answer <br> $\bullet$ OR one error made with respect to units | 2 |
| $\bullet$ Correct substitution into a correct formula | 1 |

$E_{p}=-G \frac{m_{1} m_{2}}{r}$
$E_{P}=-\frac{6.67 \times 10^{-11} \times 6.0 \times 10^{24} \times 12}{40000 \times 10^{3}}=-1.2 \times 10^{8} \mathrm{~J}$
42.

| Criteria | Marks |
| :--- | :--- |
| • Uses formula correctly | 2 |
| - Correct answer found |  |
| - Units correct | 1 |
| - Correct substitution and units |  |

43. (a)

| Marking Criteria | Marks |
| :---: | :--- |
| $\bullet$ Correct answer with units | 1 |

Mass will be the same on Alpha as it is on Earth i.e. $\frac{800}{9.8}=80 \mathrm{Kg}$ (approx..)
(b)

| Marking Criteria | Marks |
| :---: | :--- |
| $\bullet$ Correct answer | 1 |

Weight will be $\frac{3.75}{9.8} \times 800=306 N$
(c)

| Marking Criteria | Marks |
| :---: | :--- |
| $\bullet$ Correct working and answer | 2 |
| $\bullet$ Partially correct working and answer | 1 |

Use $g=\frac{G M}{r^{2}}$ for both Mars and Earth, then compare.
$r=\sqrt{\frac{G M}{g}}$
$\frac{r_{\text {mars }}}{r_{\text {earth }}}=\sqrt{ }\left(\frac{G M_{\text {mars }}}{g_{\text {mars }}} \div \frac{G M_{\text {earth }}}{g_{\text {earth }}}\right)$
$=\sqrt{\left(\frac{G M_{\text {mars }}}{g_{\text {mars }}} \times M_{\text {mars }}\right)}$
$=\sqrt{ }\left(\frac{9.8}{3.75} \times \frac{0.107358 M_{\text {earth }}}{M_{\text {earth }}}\right)$
$=\sqrt{ }\left(\frac{0.107358}{\frac{3.75}{9.8}}\right)$
$r_{\text {mars }}=0.53 r_{\text {earth }}$
OR

Use $r_{\text {mars }}=\sqrt{\frac{G M_{\text {mars }}}{g_{\text {mars }}}}$
$=\sqrt{ }\left(\frac{\left(6.67 \times 10^{-11}\right)\left(0.107358 M_{\text {earth }}\right)}{3.75}\right)$
$=3.38 \times 10^{6} \mathrm{~m}$
$r_{\text {earth }}=\sqrt{ }\left(\frac{G M_{\text {earth }}}{g_{\text {earth }}}\right)$
$=\sqrt{ }\left(\frac{\left(6.67 \times 10^{-11}\right)\left(M_{\text {earth }}\right)}{9.8}\right)$
$=6.38 \times 10^{6} \mathrm{~m}$
$r_{\text {Mars }}=0.53 r_{\text {Earth }}$
44. а) $\quad T_{T}=3 T_{R} \quad r_{T}=1.2 \times 10^{6} \mathrm{~km}$

$$
\begin{aligned}
& \frac{T_{T}^{2}}{r_{T}^{3}}=\frac{T_{R}^{2}}{r_{R}^{3}} \\
& \therefore r_{R}^{3}=\frac{T_{R}^{2}}{3\left(T_{R}^{2}\right)^{2}} \times\left(1.2 \times 10^{6}\right)^{3} \\
& \therefore r_{R}=5.77 \times 10^{8} \mathrm{~m} \text { or } 5.77 \times 10^{5} \mathrm{~km}
\end{aligned}
$$

| Criteria | Marks |
| :--- | :--- |
| Calculates correct answer with correct unit | 2 |
| Makes a single error in calculation or wrong unit | 1 |

b) $\quad \frac{r^{3}}{T^{2}}=\frac{G M}{4 \pi^{2}}$

$$
\begin{aligned}
& T=\sqrt{\frac{\left(5.77 \times 10^{8}\right)^{3} 4 \pi^{2}}{\left(6.67 \times 10^{-11}\right)\left(5.68 \times 10^{26}\right)}} \\
& T=4.47 \times 10^{5} \\
& \text { orbital velocity, } v=\frac{2 \pi r}{T} \\
& v=\frac{2 \pi\left(5.77 \times 10^{8}\right)}{4.47 \times 10^{5}} \\
& v=8103 \mathrm{~ms}^{-1} \\
& O R v=\sqrt{\frac{G M}{r}} \\
& v=\left[\frac{\left(6.67 \times 10^{-11}\right)\left(5.68 \times 10^{26}\right)}{5.77 \times 10^{8}}\right]^{\frac{1}{2}} \\
& v=8103 \mathrm{~ms}^{-1}
\end{aligned}
$$

| Criteria | Marks |
| :--- | :--- |
| Calculates correct answer using value of $r$ from (a), with correct unit | 2 |
| Makes an error/wrong unit | 1 |

$\begin{array}{lll}\text { c) } & m_{c}=3 \boldsymbol{m} & m_{z}=m \\ r_{c}=3 r & r_{z}=2 r \\ g_{c}=\frac{G m_{c}}{r_{c}^{2}} & g_{z}=\frac{G m_{2}}{r_{z}^{2}} \\ g_{c}=1 \boldsymbol{m s ^ { - 1 }} & g_{z}=\frac{G(m)}{(2 r)^{2}}\end{array}$

$$
\begin{array}{ll}
\text { i.e. } \frac{G(3 m)}{(3 r)^{2}}=1 & g_{z}=\frac{G m}{4 r^{2}} \\
\frac{G M}{3 r^{2}}=1 & \therefore g_{z}=\frac{\left(\frac{G m}{4 \pi r^{2}}\right)}{\frac{G m}{3 r^{2}}} \times 1 \\
& \therefore g_{z}=0.75 m^{-2}
\end{array}
$$

| Criteria | Marks |
| :--- | :--- |
| Calculates correct answer, with correct unit | 2 |
| Makes a single algebraic error, with correct unit | 1 |

45.a)

| Marking Criteria | Marks |
| :--- | :---: |
| Provides comprehensive definition of Newton's law of gravitation, referring to the <br> masses of both objects and the inverse of the square of the distance. | 2 |
| Provides an answer which leaves out one of the dependencies from (a) OR writes the <br> correct formula. | 1 |

The gravitational force between two objects is proportional to the product of the masses of the objects and inversely proportional to the square of the distance between them.
b)

| Marking Criteria | Marks |
| :--- | :---: |
| Accurately calculates the approximate mass of the sun. | 2 |
| Provide correct equation and substitution of values, but makes an error in the <br> calculation. | 1 |

$$
\begin{aligned}
& a=G M / r 2 \\
& M=a r 2 / G \\
& M=6 \times 103 \times(1.5 \times 1011) 2 \div 6.67 \times 10-11=2.0 \times 1036 \mathrm{~kg}
\end{aligned}
$$

46. 

| Marking Criteria | Marks |
| :--- | :---: |
| Uses 1 kg mass to determine little g | $\mathbf{2}$ |
| $g=G \frac{M_{\text {Mars }}}{R_{\text {Mars }}^{2}}=6.67 \times 10^{-11} \times \frac{6.4 \times 10^{23}}{\left(3396.2 \times 10^{3}\right)^{2}}=3.7 \mathrm{~m} / \mathrm{s}^{2}$ |  |
| One error (e.g. no unit) | $\mathbf{1}$ |

47. (a) Identifies the correct equation $\frac{R^{3}}{T^{2}}=\frac{G M}{4 \pi^{2}} \quad 1$ mark

Correctly substitutes into the equation: i.e. $M=\frac{\left(\left(6.0 \times 10^{9}\right)^{3} \times 4 \pi^{2}\right)}{\left(2.8 \times 10^{5}\right)^{2} \times 6.67 \times 10^{-11}} \quad 1$ mark
Obtains the correct answer: $1.6 \times 10^{30} \mathrm{~kg}$
1 mark
(b) Correctly substitutes into the correct equation:
i.e. $F=\frac{6,67 \times 10^{-11} \times 6.6 \times 10^{24} \times 1.63 \times 10^{20}}{\left(6.0 \times 10^{9}\right)^{2}} \quad 1$ mark

Obtains the correct answer: $2.0 \times 10^{25} \mathrm{~N}$

