

HSC Physics – Module 5: Advanced Mechanics – Motion in Gravitational Fields Short Answer Question Solutions 2

11. (a)

Criteria	Marks
Arrow used to indicate the force acting from the satellite is towards the centre of the earth	1

(b)

Criteria	Marks
Correctly calculates the initial E_p and final E_p using the appropriate radii including units (-1.209×10^{14} J and 1.164×10^{14} J respectively AND the difference between the final and initial values (4.5×10^{12} J)	3
Calculates the initial E_p and final E_p and the difference between them but uses incorrect radii (including incorrect unit conversions) OR Correctly calculates the initial E_p and final E_p using the appropriate radii including units but not the difference between them	2
Correctly calculate the initial E_p and final E_p	1

12.

(a)

Criteria	Marks
Correctly calculate the radius of the orbit of 55 cancri b (1.5×10^{10} m) and the orbital velocity (7.42×10^4 m s ⁻¹)	2
Correctly calculate the radius of orbit of 55 cancri b	1

(b)

Criteria	Marks
Correct calculation of the ratio using Kepler's law of periods (441.71:1)	2
Determines the period of 55 Cancri d (6493.21 days)	1

13.

Criteria	Marks
Uses an appropriate method to calculate a mass of 5.6×10^{26} kg	3
Correctly method of calculation in kg but with one incorrect substitution OR Correct answer with incorrect units.	2
Correctly method of calculation but with two incorrect substitutions.	1

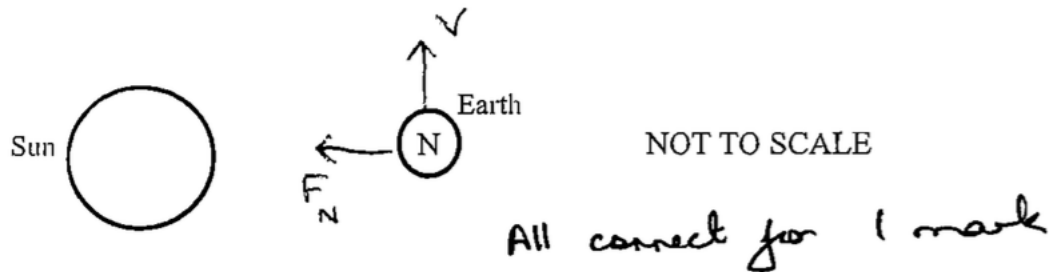
14. (a) 8.78 m/s^2 1 mark for correct number, 1 mark for correct unit

(b) $v = \left(\frac{Gm}{r} \right)^{\frac{1}{2}}$ 1 mark
 = substitution line (no carry overs paid)
 = 7690 m/s 1 mark for correct value and unit

(c) $W = E_{p \text{ final}} - E_{p \text{ initial}}$
 $= -\frac{Gmm}{r_f} - -\frac{Gmm}{r_i}$ 1 mark
 = substitution line 1 mark for correct answer and unit
 = $4.67 \times 10^9 \text{ J}$

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15. a)



(b) Show your working

$$\begin{aligned} \frac{r^3}{T^2} &= \frac{Gm}{4\pi^2} && \leftarrow 1 \text{ mark correct equation} \\ r^3 &= \frac{GmT^2}{4\pi^2} \\ &= \frac{(6.67 \times 10^{-11} \times 2 \times 10^{30} \times 9.9452 \times 10^{14})}{4\pi^2} \\ r^3 &= 3.360542 \times 10^{33} \\ r &= 1.5 \times 10^{11} \text{ m} && \leftarrow \text{max 3 sig. figs (1 mark all correct)} \\ &= 1.5 \times 10^8 \text{ km} \end{aligned}$$

16. V necessary to escape Earth's gravitational field 1 mark

$$\begin{aligned} V &= \sqrt{\frac{2GM}{r}} \\ &= \sqrt{\frac{2 \times 6.67 \times 10^{-11} \times 6 \times 10^{24}}{6.4 \times 10^6}} && 1 \text{ mark} \\ &= 11,183 \text{ ms}^{-1} && 1 \text{ mark} \end{aligned}$$

17. (a) $1.44 \times 10^6 \text{ J}$

(b) $-3.245 \times 10^6 \text{ J}$

$$\begin{aligned} \text{(c) (a) + (b)} &= \frac{-6.67 \times 10^{-11} \times 1150 \times 1.1 \times 10^{20}}{r} && (1 \text{ mark}) \\ r &= 4.67 \times 10^6 \text{ m} && (1 \text{ mark}) \end{aligned}$$

$$18. \text{(a) } g = \frac{GM}{r^2} = 0.83 \text{ m/s}^2$$

$$\begin{aligned} \text{(b) } T &= \sqrt{\frac{4\pi^2 r^3}{GM}} && (1 \text{ mark}) \\ T &= 3.24 \times 10^4 \text{ s (9 hours)} && (1 \text{ mark}) \end{aligned}$$

$$\begin{aligned} \text{(c) } \frac{GMm}{r^2} &= \frac{mv^2}{r} && \text{(or in words) (1 mark)} \\ \text{Hence } v &= \sqrt{\frac{GM}{r}} && \text{which does not depend on the mass of the satellite. (1 mark)} \end{aligned}$$

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19. (a)

Marking Criteria	Marks
<ul style="list-style-type: none"> Correct effect at both locations described 	2
<ul style="list-style-type: none"> Correct effect at either location described 	1

Sample answer:

Length of day shorter thus rotation of Earth speeds up. Centripetal acceleration of v^2/r must be subtracted from 9.8 ms^{-2} . Thus effective “g” near equator is less, say 9.7 ms^{-2} . No circular motion at south pole so no effect on value of “g”. OR Faster v gives equatorial bulge so r increases so g is less and at pole r decreases so g increases.

(b)

Marking Criteria	Marks
<ul style="list-style-type: none"> Both changes described 	2
<ul style="list-style-type: none"> One change described or both changes (period and radius) identified 	1

Sample answer:

Earth rotates faster so the satellite must also travel faster and thus it will have a shorter period to match that of the Earth. It will do so if drops to a lower orbit where “g” is greater and the speed increases to compensate.

20.

Marking Criteria	Marks
<ul style="list-style-type: none"> response provided shows evidence of proper and thorough understanding of the concepts involved, presented in a logical manner appropriate energy formulae used correctly 	4
<ul style="list-style-type: none"> sound understanding of relevant factors and relationships evident appropriate energy formulae referred to 	3
<ul style="list-style-type: none"> basic understanding of the relevant factors and relationships evident, some appropriate formulae identified 	2
<ul style="list-style-type: none"> Some understanding of a relevant concept evident 	1

Sample answer:

For a spacecraft to escape Earth’s gravitational field $PE + KE = 0$, $EK = \frac{1}{2}mv^2$ and gravitational potential energy, $EP = -Gm_1 m_2 / r$. This solves to give escape velocity = $\sqrt{2Gm/r}$. It needs to propel itself with sufficient fuel so that it can “climb up” to a potential energy of zero while still having kinetic energy and thus escape and not be drawn back by Earth’s gravity.